P01a On the contraction of protostellar clouds with different metallicities

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We examine the thermal and chemical evolution of gravitationally contracting protostellar clouds of varying metalicity (from metal-free to pop I composition gas). Before the clouds become opaque to continuum radiation, they evolve differently for different metalicity gas, namely at higher temperatures for lower metalicity clouds. However, during the adiabatic contraction phase of transient cores, the evolutionary paths of the clouds converge to a curve that is determined by fundamental physical constants only, and they coincide thereafter, regardless of their metallicity. Therefore, the physical dimension of the stellar core at the formation is the same for gas of any metalicity The mass accretion rate onto the stellar core depends on metalicity and is higher for a lower metalicity cloud, reflecting the higher temperature of the outer envelope.