S34a

IR observation and SED+spectroscopy fitting of a buried AGN in the nearby merging galaxy NGC 6240

Tamami I. Mori (The University of Tokyo), Masatoshi Imanishi (Subaru Telescope), Almudena Alonso-Herrero (Instituto de Fisica de Cantabria), Chris Packham (University of Texas at San Antonio), Cristina Ramos Almeida (Instituto de Astrofísica de Canarias), Robert Nikutta (Pontificia Universidad Católica de Chile), Omaira González-Martín (Instituto de Astrofísica de Canarias), Eric Perlman (Florida Institute of Technology), Yuriko Saito (Graduate University for Advanced Studies), & Nancy A. Levenson (Gemini Observatory)

NGC 6240 is one of the famous infrared-luminous merging galaxy with two buried AGNs in the local universe (z = 0.0245). We performed $K'(2.12 \,\mu\text{m})$, $L'(3.77 \,\mu\text{m})$, and $M'(4.68 \,\mu\text{m})$ imaging of this galaxies using IRCS on the Subaru telescope with the assistance of the AO system. We also obtained Si-2 filter band $(8.7 \,\mu\text{m})$ imaging and N-band $(7.5-13 \,\mu\text{m})$ spectroscopy with CanariCam on the Gran Telescopio Canarias (GTC). The achieved spatial resolutions of the Subaru and GTC observations were around 0.1-0.2'' and 0.4-0.5'', respectively. Combining these data with literature values, we re-evaluated the 2–30 μ m spectral energy distribution (SED) of the southern nucleus, and performed the SED+spectroscopy fitting by using CLUMPY torus models and a Bayesian fitting approach. The model fit suggests that the AGN in the southern nucleus has a high-covering factor torus and is deeply embedded in the host galaxy with an additional foreground extinction of $A_V=19$ mag, and its bolometric luminosity accounts for 37% of the whole energy of the system.